

## Science Drivers for the SPICA Mission: Extrasolar Planets and their Formation

Motohide Tamura<sup>1</sup>, Takao Nakagawa<sup>2</sup>, Hirokazu Kataza<sup>2</sup>,  
Hiroshi Shibai<sup>3</sup>, Toshio Matsumoto<sup>2</sup>, and Hideo Matsuhara<sup>2</sup>

(Email: [hide@subaru.naoj.org](mailto:hide@subaru.naoj.org))

<sup>1</sup>National Astronomical Observatory of Japan

<sup>2</sup>Institute of Space and Astronautical Science (ISAS),

Japan Aerospace Exploration Agency (JAXA),

Yoshinodai, Sagamihara, Kanagawa, Japan

<sup>3</sup>Nagoya University, Nagoya, Japan

It is needless to say that the detection and characterization of extrasolar planets is one of the most important topics of any future optical-IR observatories. The SPICA is the space mission to launch a 3.5-m diameter, cooled, single-mirror telescope working at mid- and far-infrared wavelengths. Although the spatial resolutions are not high enough to resolve the planets discovered by the radial velocity measurements, the high sensitivity of the SPICA is a powerful tool to conduct imaging and spectroscopy of possible planets and companion brown dwarfs relatively away from the central star. This potential will be enhanced if a coronagraphic instrument is equipped for SPICA, which takes advantages of its non-segmented large primary mirror. Direct observations of the formation site of such planets, protoplanetary disks, also merit from such a capability. In this contribution, we will describe the SPICA coronagraph instrument and its various scientific applications for studies on extrasolar planet, brown dwarf studies, and planet formation.